

# List of Header Columns

## I. OVERVIEW

The first table in this document lists the column headers that appear in the `.eep.track`, `.iso`, and `.iso.cmd` files, and a brief description for each. Note that not all column headers appear in each type of file. Also note that the filters in `.iso.cmd` are listed in a separate table, also found in this document. All logarithms that appear in this list are base 10. Surface and central abundances are averaged over the outer and inner  $10^{-6}\%$  of the total stellar mass.

Theoretical isochrones are provided in two flavors: basic and full. The basic isochrones contain columns such as age, stellar mass,  $\dot{M}$ ,  $\log L$ ,  $\log T_{\text{eff}}$ ,  $\log g$ , and surface and central abundances of a few elements, whereas the full isochrones are much more comprehensive. Columns that appear in the basic file are marked by an asterisk (\*) in the table below.

The second table in this document lists the primary equivalent evolutionary points (EEPs) and their corresponding EEP number.

The third table in this document lists the currently available filters. This is only an initial set and will expand over time. The zero points depend on the photometric system—check the file headers for details.

TABLE I: EEP Track and Isochrone Column Headers

Column Name	Description
<i>Appears in .track.eep Only</i>	
star_age	Age in years
<i>Appears in .iso.cmd Only</i>	
Zsurf	Surface metal mass fraction
<i>Appears in .iso and .iso.cmd Only</i>	
EEP*	Equivalent Evolutionary Point number
initial_mass*	Initial mass in $M_{\odot}$
log10_isochrone_age_yr*	Age of the isochrone in log years
OR	
isochrone_age_yr*	Age of the isochrone in years
<i>Appears in .track.eep, .iso, and .iso.cmd</i>	
star_mass*	Current mass in $M_{\odot}$
star_mdot*	Mass loss rate in $M_{\odot}/\text{year}$
he_core_mass*	Mass of the helium-rich core in $M_{\odot}$
c_core_mass*	Mass of the carbon-rich core in $M_{\odot}$
o_core_mass	Mass of the oxygen-rich core in $M_{\odot}$
log_L*	Log bolometric luminosity in $L_{\odot}$
log_L.div.Ledd	Log ratio of bolometric luminosity and Eddington luminosity, where the Eddington luminosity is a mass-weighted average over the optical depth $\tau$ between 1 and 100
log_LH*	Log hydrogen-burning luminosity in $L_{\odot}$
log_LHe*	Log helium-burning luminosity in $L_{\odot}$
log_LZ	Log total burning luminosity excluding H-burn, He-burn, and photodisintegrations in $L_{\odot}$
log_Teff*	Log effective temperature in K
log_abs_Lgrav	Log gravitational potential luminosity in $L_{\odot}$
log_R*	Log radius in $R_{\odot}$
log_g*	Log surface gravity in $\text{cm s}^{-2}$
log_surf_cell_z	Log surface mass fraction in metals
(previously named log_surf_z)	
surf_avg_omega	Surface angular rotation speed
surf_avg_v_rot	Surface rotation speed
surf_num_c12_div_num_o16	Ratio of surface number densities of $^{12}\text{C}$ and $^{16}\text{O}$
v_wind_Km_per_s	Wind speed $v_w \equiv \kappa \dot{M} / 4\pi R \tau$ , where $\kappa \equiv$ opacity and $\tau = 2/3$ , in km/s
surf_avg_omega_crit	Surface (mass-averaged down to $\tau = 100$ ) critical angular rotation speed
surf_avg_omega_div_omega_crit	Ratio of surface and critical angular rotation speeds
surf_avg_v_crit	Surface critical/breakup rotation speed
surf_avg_v_div_v_crit	Ratio of surface and critical rotation speeds
surf_avg_Lrad_div_Ledd	Ratio of surface radiative luminosity and Eddington luminosity
v_div_csound_surf	Ratio of velocity and sound speed at the surface
surf_r_equatorial_div_r	Equatorial radius in units of “effective radius” computed following Endal & Sofia 1976
surf_r_polar_div_r	Polar radius in units of “effective radius”

total_angular_momentum	computed following Endal & Sofia 1976
surface_h1*	Total angular momentum in the stellar interior
surface_he3*	Surface mass fraction in $^1\text{H}$
surface_he4*	Surface mass fraction in $^3\text{He}$
surface_li7	Surface mass fraction in $^4\text{He}$
surface_be9	Surface mass fraction in $^7\text{Li}$
surface_b11	Surface mass fraction in $^9\text{Be}$
surface_c12*	Surface mass fraction in $^{11}\text{B}$
surface_c13	Surface mass fraction in $^{12}\text{C}$
surface_n14	Surface mass fraction in $^{13}\text{C}$
surface_o16*	Surface mass fraction in $^{14}\text{N}$
surface_f19	Surface mass fraction in $^{16}\text{O}$
surface_ne20	Surface mass fraction in $^{19}\text{F}$
surface_na23	Surface mass fraction in $^{20}\text{Ne}$
surface_mg24	Surface mass fraction in $^{23}\text{Na}$
surface_si28	Surface mass fraction in $^{24}\text{Mg}$
surface_s32	Surface mass fraction in $^{28}\text{Si}$
surface_ca40	Surface mass fraction in $^{32}\text{S}$
surface_ti48	Surface mass fraction in $^{40}\text{Ca}$
surface_fe56	Surface mass fraction in $^{48}\text{Ti}$
log_center_T*	Surface mass fraction in $^{56}\text{Fe}$
log_center_Rho*	Log central temperature in K
center_degeneracy	Log central density in $\text{g cm}^{-3}$
	Central electron chemical potential in $k_b T$ ,
	where $k_b \equiv$ Boltzmann constant and $T \equiv$ temperature
center_omega	Central angular rotation speed
center_gamma*	Central plasma interaction parameter
	$\bar{Z}^2 e^2 / a_i k_b T$ , where $\bar{Z} \equiv$ average ion charge,
	$e \equiv$ electron charge, and $a_i \equiv$ mean ion spacing
mass_conv_core	Mass of the convective core in $M_\odot$
center_h1*	Center mass fraction in $^1\text{H}$
center_he4*	Center mass fraction in $^4\text{He}$
center_c12*	Center mass fraction in $^{12}\text{C}$
center_n14	Center mass fraction in $^{14}\text{N}$
center_o16	Center mass fraction in $^{16}\text{O}$
center_ne20	Center mass fraction in $^{20}\text{Ne}$
center_mg24	Center mass fraction in $^{24}\text{Mg}$
center_si28	Center mass fraction in $^{28}\text{Si}$
pp	Log luminosity from pp-chain
cno	Log luminosity from CNO-cycle
tri_alfa	Log luminosity from triple $\alpha$
burn_c	Log luminosity from carbon-burning
burn_n	Log luminosity from nitrogen-burning
burn_o	Log luminosity from oxygen-burning
c12_c12	Log luminosity from carbon-carbon burning
delta_nu	Large frequency separation for p-modes in $\mu\text{Hz}$
delta_Pg	Period spacing for $l = 1$ g-mode in seconds
nu_max	Frequency of maximum power in $\mu\text{Hz}$
	as estimated from scaling relations
acoustic_cutoff	Maximum frequency for p-modes at surface
max_conv_vel_div_sound	Maximum ratio of convective velocity
	and sound speed in the stellar interior
max_gradT_div_grada	Maximum ratio of $\nabla_T$ and $\nabla_{\text{ad}}$ in the stellar interior
gradT_excess_alpha	Denoted by $\alpha_\nabla$ and referred to as the ‘‘Smoothing
	parameter for $\text{MLT}_{++}$ ’’ in Paxton et al. 2013.

<code>min.Pgas_div_P</code>	Number between 0 and 1 describing the effectiveness with which the $MLT_{++}$ prescription is used to aid the evolution calculations by reducing the superadiabaticity
<code>max.Lrad_div_Ledd</code>	Minimum ratio of gas pressure to the total pressure in the stellar interior
<code>e_thermal</code>	Maximum ratio of radiative luminosity and Eddington luminosity in the interior
<code>envelope_binding_energy</code>	Total thermal energy in the stellar interior in ergs
<code>conv_env_top_mass</code>	Total binding energy in the envelope, defined as the H-rich region, i.e., hydrogen mass fraction is above $10^{-4}$
<code>conv_env_bot_mass</code>	Location of the top of the convection zone in mass
<code>conv_env_top_radius</code>	Location of the bottom of the convection zone in mass
<code>conv_env_bot_radius</code>	Location of the top of the convection zone in radius
<code>conv_env_turnover_time_lt</code>	Location of the bottom of the convection zone in radius
<code>conv_env_turnover_time_lb</code>	Turnover time (seconds) in the convective envelope, defined to be the mixing length ( $\alpha_{MLT}H_P$ ) divided by the $v_{conv}$ evaluated one full mixing length above <code>conv_env_bot_radius</code>
<code>conv_env_turnover_time_g</code>	Turnover time (seconds) in the convective envelope, defined to be the mixing length ( $\alpha_{MLT}H_P$ ) divided by the $v_{conv}$ evaluated one half mixing length above <code>conv_env_bot_radius</code>
<code>phase*</code>	“Globally-averaged” turnover time (seconds) in the convective envelope, $dr/v_{conv}(r)$ integrated over the entire convective envelope
	FSPS phase type defined as follows: -1=PMS, 0=MS, 2=RGB, 3=CHeB, 4=EAGB, 5=TPAGB, 6=postAGB, 9=WR Caution: There may be overlap between MS and WR for very massive stars. Always double-check!

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TABLE II: Primary EEPs

Primary EEP	EEP Number <sup>a</sup>	Phase
1	1	pre-main sequence (PMS)
2	202	zero age main sequence (ZAMS)
3	353	intermediate age main sequence (IAMS)
4	454	terminal age main sequence (TAMS)
5	605	tip of the red giant branch (RGBTip)
6	631	zero age core helium burning (ZACHeB) <sup>b</sup>
7	707	terminal age core helium burning (TACHeB) <sup>c</sup>
Low Mass Type		
8	808	thermally pulsating asymptotic giant branch (TPAGB)
9	1409	post asymptotic giant branch (post-AGB)
10	1710	white dwarf cooling sequence (WDCS)
High Mass Type		
8	808	carbon burning (C-burn)

<sup>a</sup>Also equivalent to  $i + 1$  where  $i$  is the index of the array (zero-based) containing the evolutionary track.

<sup>b</sup>i.e., zero age horizontal branch; ZAHB for low-mass stars.

<sup>c</sup>terminal age horizontal branch; TAHB.

TABLE III: Currently Available Filters

Name	Reference
Bessell_U	[1]
Bessell_B	
Bessell_V	
Bessell_R	
Bessell_I	
2MASS_J	[2]
2MASS_H	
2MASS_Ks	
SDSS_u	[3]
SDSS_g	
SDSS_r	
SDSS_i	
SDSS_z	
WFPC2_F218W	[4]
WFPC2_F255W	
WFPC2_F300W	
WFPC2_F336W	
WFPC2_F439W	
WFPC2_F450W	
WFPC2_F555W	
WFPC2_F606W	
WFPC2_F622W	
WFPC2_F675W	

WFPC2\_F791W  
WFPC2\_F814W  
WFPC2\_F850LP

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ACS\_HRC\_F220W [5]  
ACS\_HRC\_F250W  
ACS\_HRC\_F330W  
ACS\_HRC\_F344N  
ACS\_HRC\_F435W  
ACS\_HRC\_F475W  
ACS\_HRC\_F502N  
ACS\_HRC\_F550M  
ACS\_HRC\_F555W  
ACS\_HRC\_F606W  
ACS\_HRC\_F625W  
ACS\_HRC\_F658N  
ACS\_HRC\_F660N  
ACS\_HRC\_F775W  
ACS\_HRC\_F814W  
ACS\_HRC\_F850LP  
ACS\_HRC\_F892N  
ACS\_WFC\_F435W  
ACS\_WFC\_F475W  
ACS\_WFC\_F502N  
ACS\_WFC\_F550M  
ACS\_WFC\_F555W  
ACS\_WFC\_F606W  
ACS\_WFC\_F625W  
ACS\_WFC\_F658N  
ACS\_WFC\_F660N  
ACS\_WFC\_F775W  
ACS\_WFC\_F814W  
ACS\_WFC\_F850LP  
ACS\_WFC\_F892N

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WFC3\_UVIS\_F200LP [6]  
WFC3\_UVIS\_F218W  
WFC3\_UVIS\_F225W  
WFC3\_UVIS\_F275W  
WFC3\_UVIS\_F280N  
WFC3\_UVIS\_F300X  
WFC3\_UVIS\_F336W  
WFC3\_UVIS\_F343N  
WFC3\_UVIS\_F350LP  
WFC3\_UVIS\_F373N  
WFC3\_UVIS\_F390M  
WFC3\_UVIS\_F390W  
WFC3\_UVIS\_F395N  
WFC3\_UVIS\_F410M  
WFC3\_UVIS\_F438W  
WFC3\_UVIS\_F467M  
WFC3\_UVIS\_F469N  
WFC3\_UVIS\_F475W  
WFC3\_UVIS\_F475X  
WFC3\_UVIS\_F487N  
WFC3\_UVIS\_F502N  
WFC3\_UVIS\_F547M

WFC3\_UVIS\_F555W  
 WFC3\_UVIS\_F600LP  
 WFC3\_UVIS\_F606W  
 WFC3\_UVIS\_F621M  
 WFC3\_UVIS\_F625W  
 WFC3\_UVIS\_F631N  
 WFC3\_UVIS\_F645N  
 WFC3\_UVIS\_F656N  
 WFC3\_UVIS\_F657N  
 WFC3\_UVIS\_F658N  
 WFC3\_UVIS\_F665N  
 WFC3\_UVIS\_F673N  
 WFC3\_UVIS\_F680N  
 WFC3\_UVIS\_F689M  
 WFC3\_UVIS\_F763M  
 WFC3\_UVIS\_F775W  
 WFC3\_UVIS\_F814W  
 WFC3\_UVIS\_F845M  
 WFC3\_UVIS\_F850LP  
 WFC3\_UVIS\_F953N  
 WFC3\_IR\_F098M  
 WFC3\_IR\_F105W  
 WFC3\_IR\_F110W  
 WFC3\_IR\_F125W  
 WFC3\_IR\_F126N  
 WFC3\_IR\_F127M  
 WFC3\_IR\_F128N  
 WFC3\_IR\_F130N  
 WFC3\_IR\_F132N  
 WFC3\_IR\_F139M  
 WFC3\_IR\_F140W  
 WFC3\_IR\_F153M  
 WFC3\_IR\_F160W  
 WFC3\_IR\_F164N  
 WFC3\_IR\_F167N

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IRAC\_3.6 [7]  
 IRAC\_4.5  
 IRAC\_5.8  
 IRAC\_8.0

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UKIDSS\_Z [8]  
 UKIDSS\_Y  
 UKIDSS\_J  
 UKIDSS\_H  
 UKIDSS\_K

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CFHT\_u [9]  
 CFHT\_g  
 CFHT\_r  
 CFHT\_i\_new  
 CFHT\_i\_old  
 CFHT\_z

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WISE\_W1 [10]  
 WISE\_W2  
 WISE\_W3  
 WISE\_W4

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Strömgren_u	[11]
Strömgren_v	
Strömgren_b	
Strömgren_y	
PS_g	[12]
PS_r	
PS_i	
PS_z	
PS_y	
PS_w	
PS_open	
GALEX_FUV	[13]
GALEX_NUV	
DECam_u	[14]
DECam_g	
DECam_r	
DECam_i	
DECam_z	
DECam_Y	
SkyMapper_u	[15]
SkyMapper_v	
SkyMapper_g	
SkyMapper_r	
SkyMapper_i	
SkyMapper_z	
Washington_C	[16]
Washington_M	
Washington_T1	
Washington_T2	
DD051_vac	[17]
DD051_f31	
Kepler_Kp	[18]
Kepler_D51	
LSST_u	[19]
LSST_g	
LSST_r	
LSST_i	
LSST_z	
LSST_y	
JWST_F070W	[20]
JWST_F090W	
JWST_F115W	
JWST_F140M	
JWST_F150W2	
JWST_F150W	
JWST_F162M	
JWST_F164N	
JWST_F182M	
JWST_F187N	
JWST_F200W	
JWST_F210M	
JWST_F212N	



JWST_F250M	
JWST_F277W	
JWST_F300M	
JWST_F322W2	
JWST_F323N	
JWST_F335M	
JWST_F356W	
JWST_F360M	
JWST_F405N	
JWST_F410M	
JWST_F430M	
JWST_F444W	
JWST_F460M	
JWST_F466N	
JWST_F470N	
JWST_F480M	
<hr/>	
Swift_UVW2	[21]
Swift_UVM2	
Swift_UVW1	
Swift_U	
Swift_B	
Swift_V	
<hr/>	
Hipparcos_Hp	[22]
<hr/>	
Tycho_B	[23]
Tycho_V	
<hr/>	
Gaia_G_DR2Rev	[24]
Gaia_BP_DR2Rev	
Gaia_RP_DR2Rev	
<hr/>	
TESS	[25]
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- [4] Holtzman et al. (1995)
- [5] [www.stsci.edu/hst/acs/analysis/throughputs](http://www.stsci.edu/hst/acs/analysis/throughputs)
- [6] [www.stsci.edu/hst/wfc3/ins\\_performance/filters/](http://www.stsci.edu/hst/wfc3/ins_performance/filters/)
- [7] Fazio et al. (2004)
- [8] Hewett et al. (2006)
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- [11] Bessell (2011)
- [12] Tonry et al. (2012)
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- [23] Bessell & Murphy (2012)
- [24] [https://www.cosmos.esa.int/web/gaia/iow\\_20180316](https://www.cosmos.esa.int/web/gaia/iow_20180316)
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